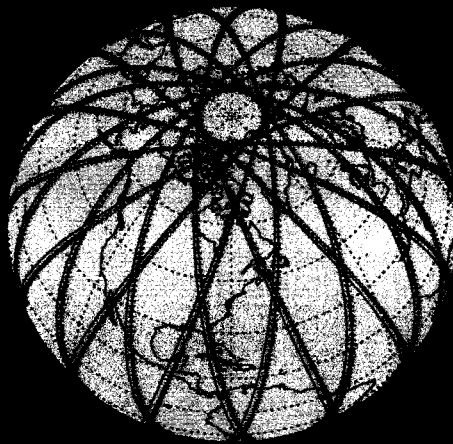


**A summary presentation emphasizing  
the EOS MLS Scientific Objectives,  
Measurements and Data Products**

**An updated copy of this presentation, and more information on the MLS experiments can be found on the MLS web site: <http://mls.jpl.nasa.gov>.**

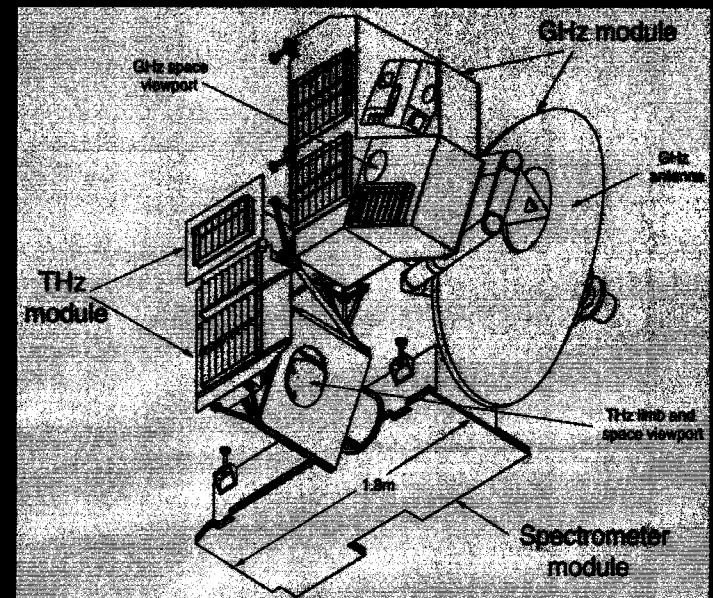
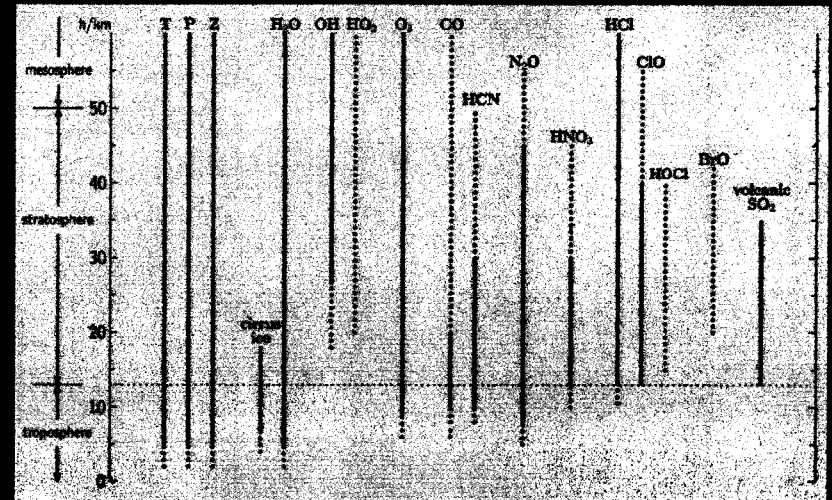
***An Overview of the EOS MLS Experiment*, available on the web site,  
is good starting point for more information on EOS MLS.**

- Determining if stratospheric ozone is recovering as expected
- Improving knowledge of processes that affect climate variability
- Helping understand ozone pollution in the upper troposphere



- Passive observations of thermal emission from the atmospheric limb
- Made day and night

- Millimeter and submillimeter wavelength heterodyne radiometers in 5 broad bands:
  - 118 GHz, primarily for temperature and pressure
  - 190 GHz, primarily for  $\text{H}_2\text{O}$  and  $\text{HNO}_3$
  - 240 GHz, primarily for  $\text{O}_3$  and  $\text{CO}$
  - 640 GHz, primarily for  $\text{HCl}$ ,  $\text{ClO}$ ,  $\text{BrO}$ ,  $\text{HO}_2$ ,  $\text{N}_2\text{O}$
  - 2.5 THz, primarily for  $\text{OH}$
- On NASA's EOS CHEMISTRY ('CHEM') satellite, to start observations in 2003 with at least 5 years operational lifetime



will be used to monitor the recovery of the ozone layer.

With the addition of the Aura satellite, the Earth Science Enterprise will have a comprehensive suite of instruments to help understand atmospheric global change.

The Aura satellite will help scientists understand the following:

• **Global climate change**

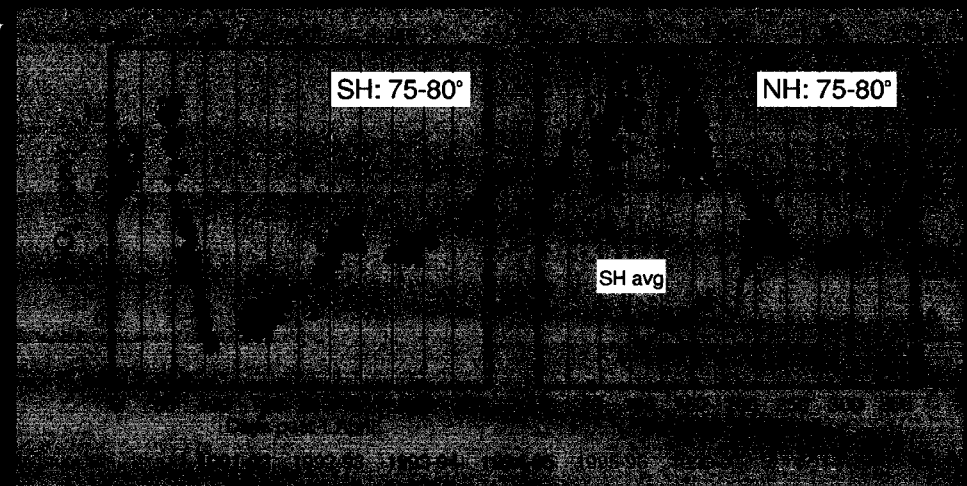
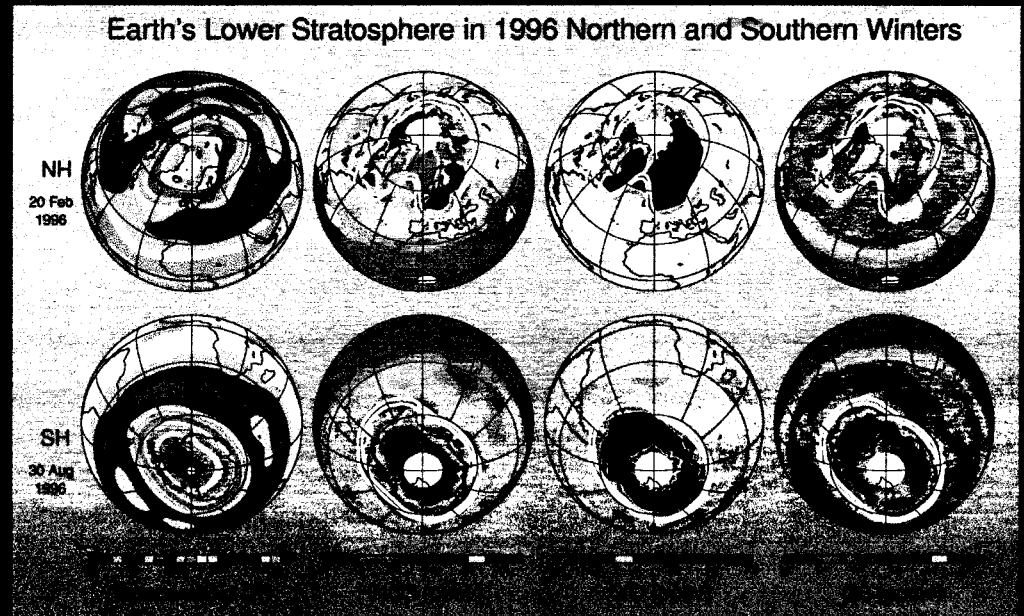
- **Determining if stratospheric ozone is recovering as expected**
- **Improving knowledge of processes that affect climate variability**
- **Helping understand ozone pollution in the upper troposphere**
- **Together these instruments provide a comprehensive suite of information for helping understand atmospheric global change.**

## EOS MLS Scientific Objective

# Determining if Stratospheric Ozone is Recovering as Expected

- Are stratospheric chlorine, ozone chemistry, and ozone responding to regulations as expected?
- Will ozone recovery be delayed by climate changes?
  - cooling of lower stratosphere
  - increases in stratospheric  $H_2O$
  - changes in atmospheric circulation
- Is Arctic on the verge of severe denitrification that could lead to increased ozone depletion?
  - due to changing climate
- Do we adequately understand chemistry and transport in the stratosphere at all altitudes and latitudes?
- How will volcanoes affect recovery?

UARS MLS  $HO_x$ ,  $ClO_x$ , and  $O_3$

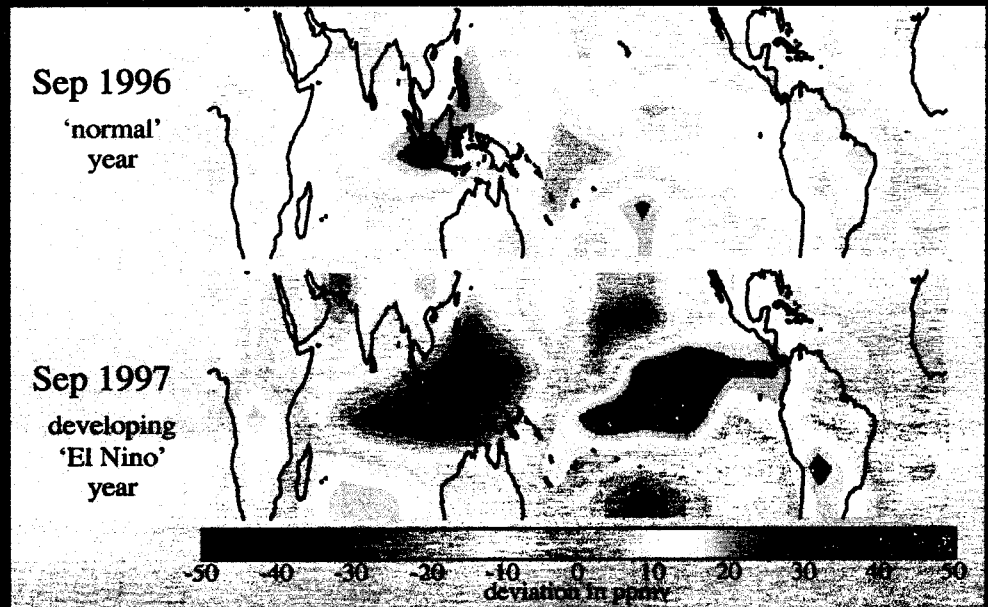


## EOS MLS Scientific Objective

# Improving Knowledge of Processes Affecting Climate Variability

### UARS MLS Upper Trop H<sub>2</sub>O Variations and El Niño

- How do feedback mechanisms involving upper tropospheric H<sub>2</sub>O affect climate variability?
- What are the atmospheric processes that control upper tropospheric H<sub>2</sub>O abundances?
- How do variations in sea surface temperature affect upper trop H<sub>2</sub>O (and thus climate)?
- How do lower stratospheric H<sub>2</sub>O and O<sub>3</sub>, and possibly N<sub>2</sub>O and Arctic vortex variations, affect climate?

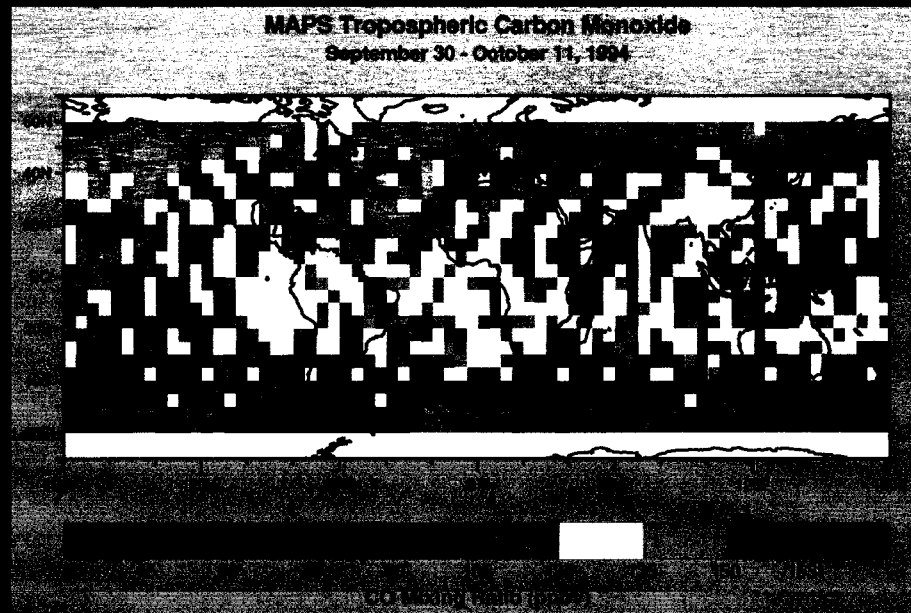
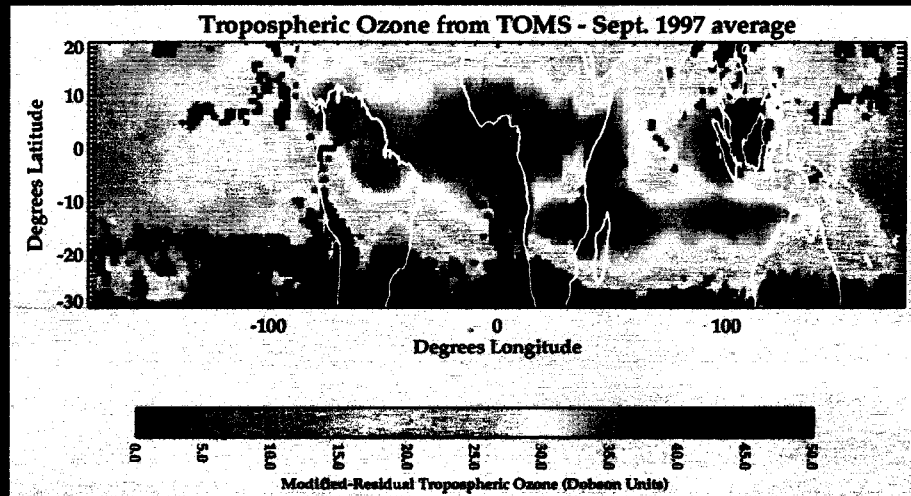


## EOS MLS Scientific Objective

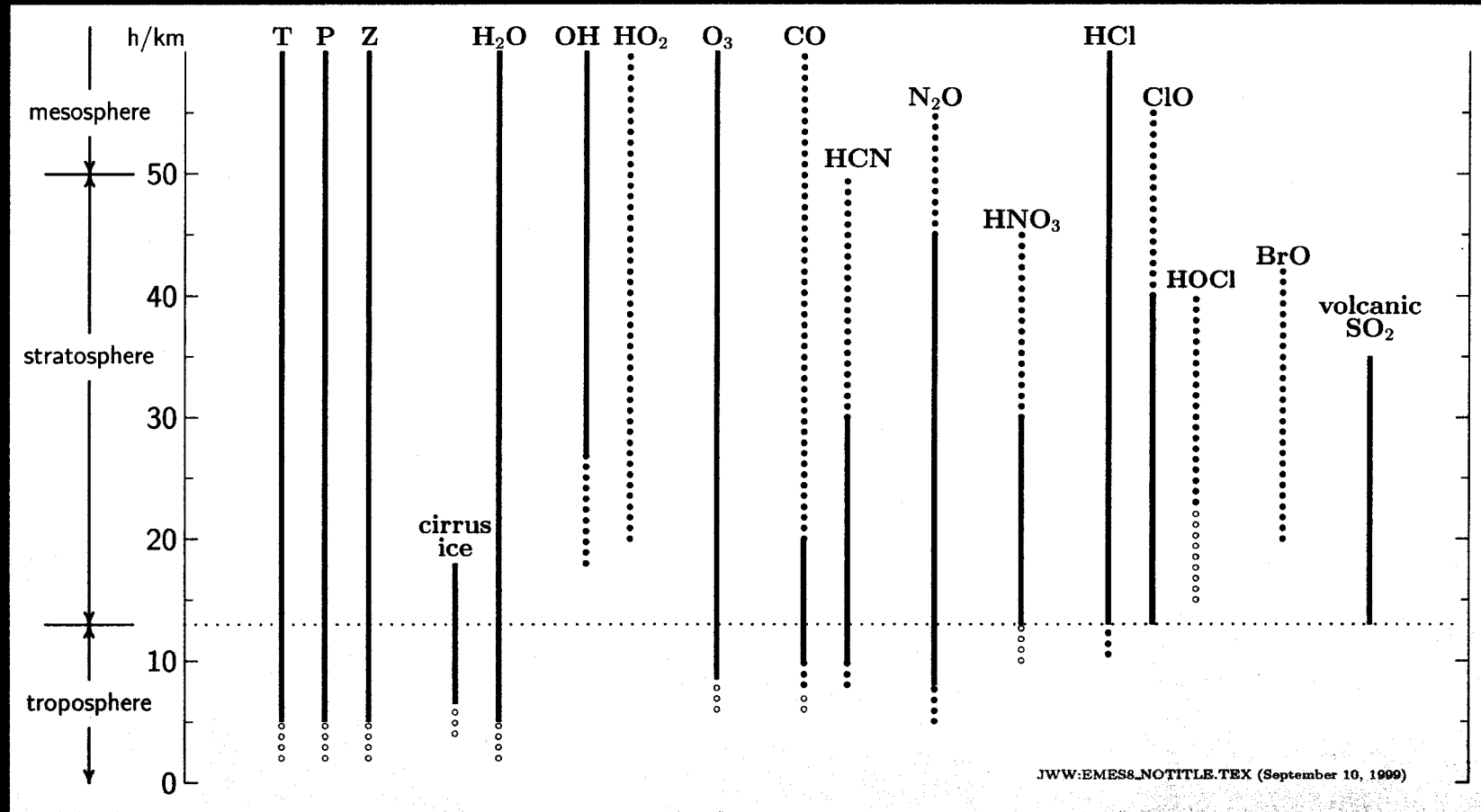
# Helping Understand Ozone Pollution in Upper Troposphere

TOMS residual tropospheric  $O_3$  from Anne Thompson, GSFC  
MAPS tropospheric CO from Vickie Connors, JPLRC

- What is the global distribution and variation of  $O_3$  in the upper troposphere?
- What are the dominant sources of upper tropospheric  $O_3$ ?
- How is regional pollution related to global upper tropospheric  $O_3$  pollution?
- How might expected increases in upper tropospheric  $O_3$  affect global air quality?



# EOS MLS Measurements Summary



**Solid lines indicate useful individual profiles or daily global maps (Z is geopotential height).**

**Dotted lines indicate zonal (or other) averages are expected to be required for useful precision.**

**Open circles indicate goals for more difficult measurements.**

**Measurements of T, P, Z, H<sub>2</sub>O, O<sub>3</sub>, HCl, OH and CO extend higher than the 60 km shown here. CH<sub>3</sub>CN, not yet shown in above plot, will also be measured - based on recent UARS MLS results.**

# EOS MLS Geophysical Data Products and Files

- **Files contain value and precision for each datum**
  - Also include diagnostic quantities and ancillary information such as: time, lat, lon, local solar time, local solar zenith angle, line-of-sight angle
  - Vertical coordinate is  $\log_{10}(P)$
- **Level 2 files**
  - Contain individual retrievals, equally-spaced along orbit for normal operation
  - A separate daily file for each standard data product (~2 Mbyte each)
  - Analogous to UARS 'L3AT' files
- **Level 3 files**
  - Daily maps: separate daily file for each standard data product (~3 Mb each)
  - Monthly maps: monthly file with all std products together (~50 Mb each)
  - Daily / monthly zonal means: daily and monthly files with all standard products grouped together (~3 Mb each)

## MLS Standard Geophysical Data Products

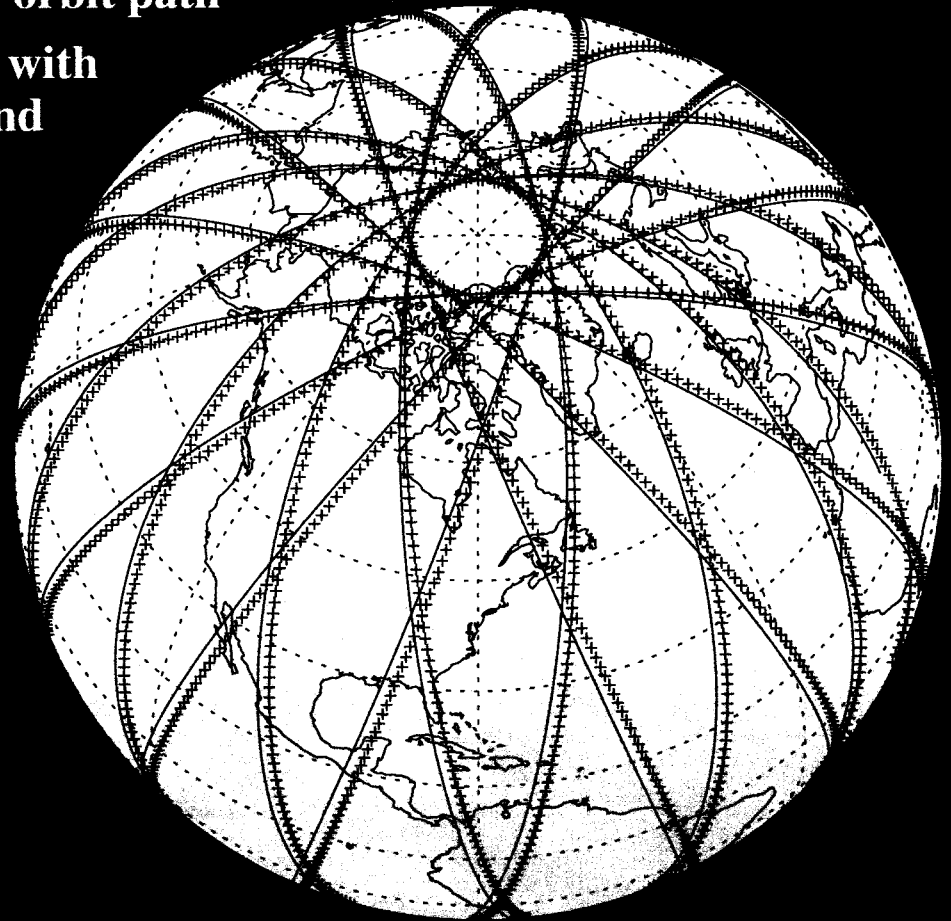
data product (alphabetical order)	~ useful vertical range			
	pressure / hPa		height / km	
	max	min	min	max
BrO mixing ratio	50	1	20	50
cirrus ice average density	500	50	5	20
ClO mixing ratio	100	0.1	15	65
CO mixing ratio	300	0.001	8	95
geopotential height	500	0.01	5	80
H <sub>2</sub> O mixing ratio	500	0.001	5	95
HCl mixing ratio	100	0.01	15	80
HCN mixing ratio	500	1	5	50
HNO <sub>3</sub> mixing ratio	100	1	15	50
HO <sub>2</sub> mixing ratio	50	0.1	20	65
HOCl mixing ratio	25	1	25	50
N <sub>2</sub> O mixing ratio	500	1	5	50
O <sub>3</sub> mixing ratio	300	0.001	8	95
O <sub>3</sub> column in stratosphere	(not applicable)		(not applicable)	
OH mixing ratio	50	0.01	20	80
relative humidity w.r.t. ice	500	50	5	20
SO <sub>2</sub> mixing ratio	100	3	15	40
temperature	500	0.001	5	95

**Note:** CH<sub>3</sub>CN will probably be added as a standard product: recent results from UARS MLS indicate that it can be measured by EOS MLS.

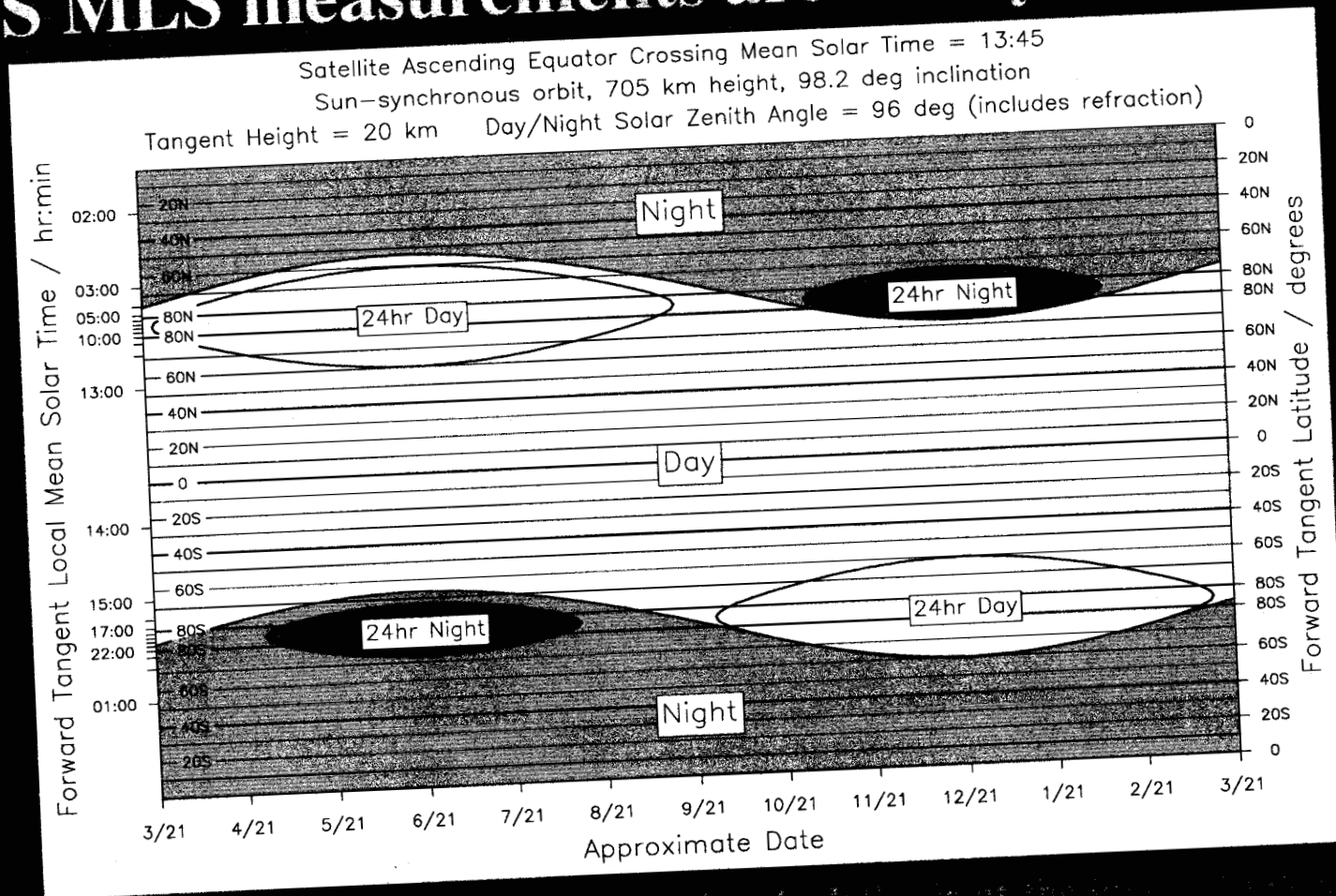


# EOS MLS Measurements

- All measurements made continuously, day and night
  - but, if needed to reduce power, can time-share
- Limb scan & calibration each  $1.5^\circ$  along track (each 25s, 165 km)
  - MLS views in forward direction along orbit path
  - Limb scans are synchronized to orbit, with 240 scans each orbit; scans over EQ and at symmetric latitudes in NH and SH
- Each day's measurement coverage is shown at right
  - each cross gives the location of a limb scan and retrieved profile (line shows subsatellite path)
  - the tangent point locus during each limb scan is nearly vertical
  - Same SH coverage as NH shown here
- Resolution vs. precision trade-offs



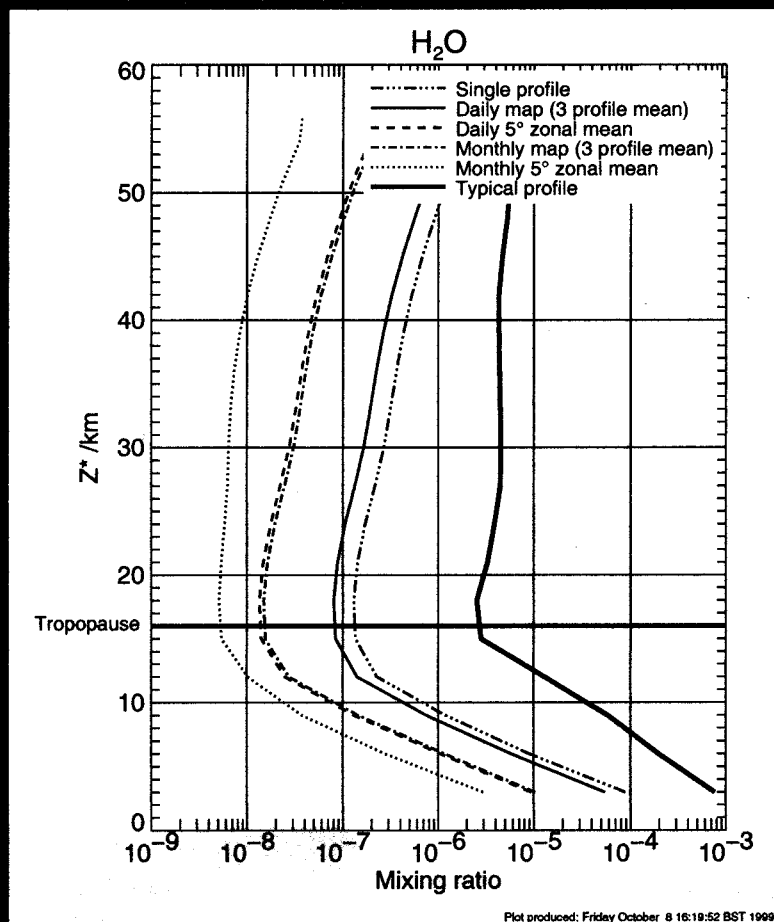
# Annual variation of latitude range where EOS MLS measurements are in day and in night



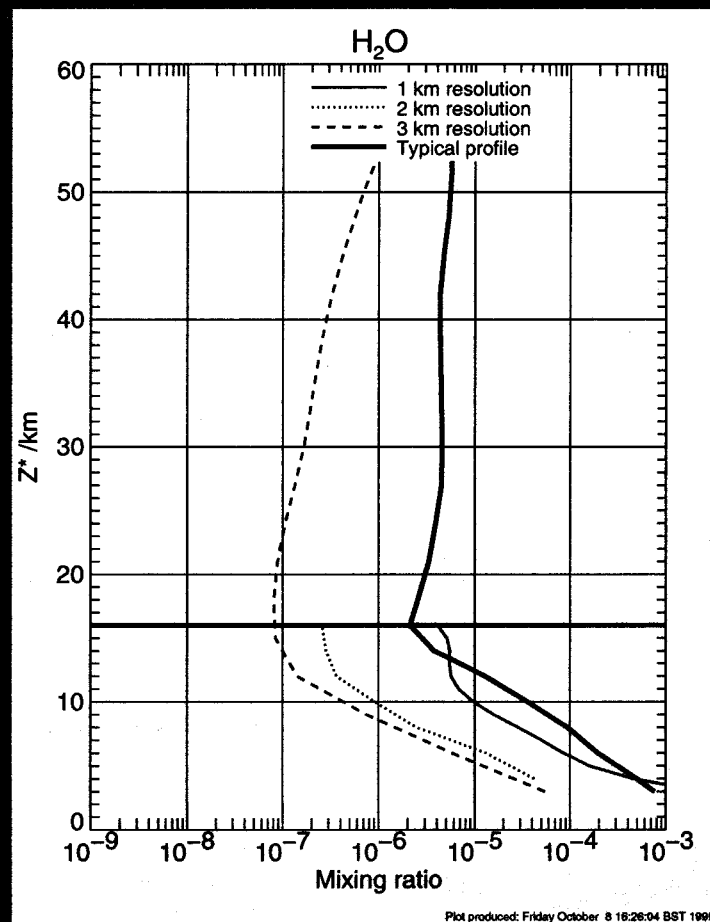
The right vertical axis and horizontal lines give the latitude of the forward tangent point where EOS MLS makes measurements; the left axis gives the local mean solar time. The shaded area is in the direction, from the equator, where the satellite is in the night.

# EOS MLS Measurement Precision Example: tropical H<sub>2</sub>O

Various H<sub>2</sub>O products at  
3 km vertical resolution



Daily maps at 3, 2, 1 km  
vertical resolution in troposphere  
(and 3 km in stratosphere)



# Some features of MLS technique

- **MLS can 'see through' dense aerosol and cirrus**
  - Clouds/aerosol have much smaller effect than at IR, visible, UV wavelengths
  - Some cirrus signals, although optically thin, are measurable by MLS and provide unique ice information
- **Spectral lines are resolved with several channels**
  - ⇒ **robust data interpretation throughout stratosphere, mesosphere, and most of upper troposphere**
  - Continuum emission from H<sub>2</sub>O (vapor, ice, liquid) and dry air limits this aspect of robustness at lowest altitudes
- **Accurate and stable calibration**
  - Upper limit of 0.02% change in UARS MLS antenna reflectivity over 5 years in orbit - this is thought to be dominant contributor to overall calibration degradation, and similar performance expected from EOS MLS
  - Composition measurements can be 'self-calibrated' (to ~1%) from temperature
- **Heritage from UARS MLS (and balloon MLS, balloon LIDAR)**
  - More than 180 MLS-related peer-reviewed scientific publications to date (updated list available on MLS web site, <http://mls.jpl.nasa.gov>)
  - EOS MLS has broader spectral coverage & finer spatial sampling than UARS

# The EOS MLS Instrument

- **Heterodyne radiometers operating in 5 broad mm/submm bands**
  - 118 GHz radiometer: primarily for temperature and tangent pressure reference
  - 190 GHz radiometer: primarily for  $\text{H}_2\text{O}$  and  $\text{HNO}_3$
  - 240 GHz radiometer: primarily for  $\text{O}_3$  and  $\text{CO}$
  - 640 GHz radiometer: primarily for  $\text{HCl}$ ,  $\text{ClO}$ ,  $\text{BrO}$ ,  $\text{HO}_2$  and  $\text{N}_2\text{O}$
  - 2.5 THz radiometer: primarily for  $\text{OH}$
- **Advanced technology**
  - Planar submm mixers
  - Integrated circuits
  - Composite materials
  - Compact gas laser for THz local oscillator and solid-state sources for GHz local oscillators
- **JPL designed and developed**
  - with many subsystems procured from industry
- **Some overall characteristics**
  - ~450 kg, ~550 W, 100 kb/s
  - 5-year on-orbit design lifetime
  - Ambient temperature operation

